

## Direct Numerical Simulation of Pluto's Extended Atmosphere

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The structure and escape of Pluto's upper atmosphere is modeled using the Direct Simulation Monte Carlo (DSMC) method. Several atmospheric compositions are studied consisting primarily of N<sub>2</sub> with differing mixing fractions of CH<sub>4</sub> and CO. Radial profiles of temperature, density and species distributions are calculated in a spherically symmetric domain with a varying gravitational body force. Lower boundary conditions are obtained from a hydrostatic, radiative model of Pluto's lower atmosphere (Strobel *et al.* 1996) and results for pure methane are compared against Trafton's hydrodynamic, extended atmosphere model (Trafton *et al.* 1987, 1989 and Clarke *et al.* 1992). Using the DSMC approach the atmosphere is modeled statistically by extrapolating from the motions and collisions of a relatively small number of representative molecules. This method can solve fully viscous, compressible, unsteady flow problems and allows for rarefied and non-LTE effects and the diffusive separation of multi-species gases.

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